

# Hunting for dark photons from Higgs boson decays with the ATLAS detector: a data-driven approach to the estimation of backgrounds in events with a photon and missing transverse momentum

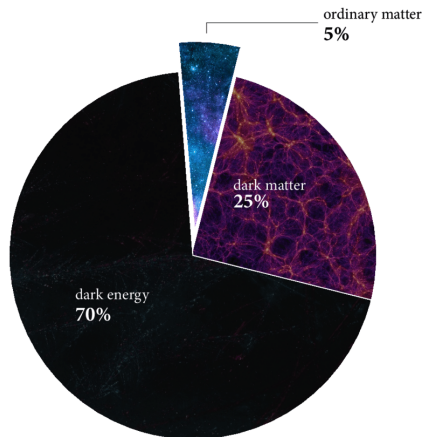
Tesi di Laurea Magistrale in Fisica di: Giulia Maineri

Relatori: Prof. Marcello Fanti, Dott.ssa Silvia Resconi, Dott.ssa Federica Piazza



UNIVERSITÀ DEGLI STUDI DI MILANO  
FACOLTÀ DI SCIENZE E TECNOLOGIE

# Dark Matter



## Dark Sector

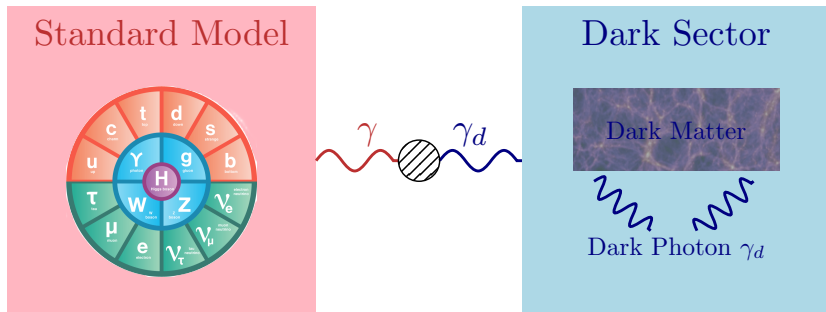


Dark Matter

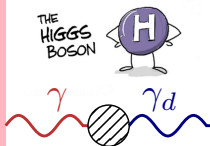
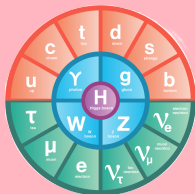


Dark Photon  $\gamma_d$

# A portal to the dark sector



## Standard Model



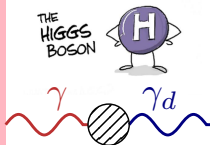
## Dark Sector



Dark Matter

Dark Photon  $\gamma_d$

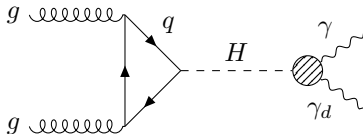
## Standard Model



## Dark Sector

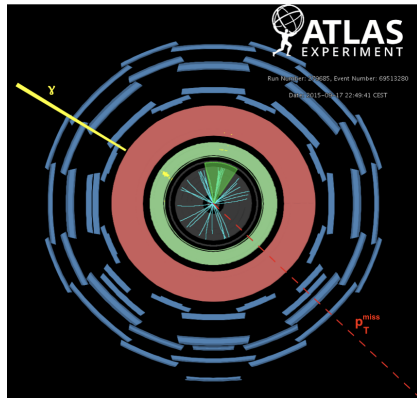
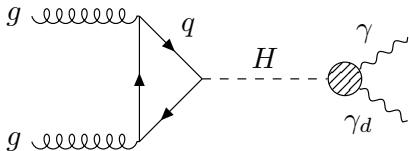


Dark Photon  $\gamma_d$

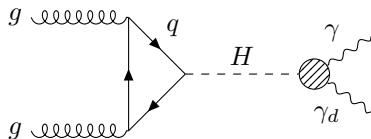


**Signal:**  $gg \rightarrow H \rightarrow \gamma\gamma_d$

# Missing transverse momentum



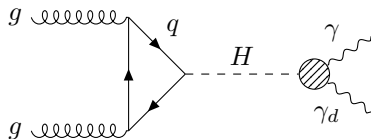
$$\vec{p}_T^{miss} = - \left[ \sum_e \vec{p}_T^{(e)} + \sum_\mu \vec{p}_T^{(\mu)} + \sum_\gamma \vec{p}_T^{(\gamma)} + \sum_\tau \vec{p}_T^{(\tau)} + \sum_{jet} \vec{p}_T^{(jet)} + \sum_{soft} \vec{p}_T^{(soft)} \right]$$



**Signal:**  $gg \rightarrow H \rightarrow \gamma\gamma_d$

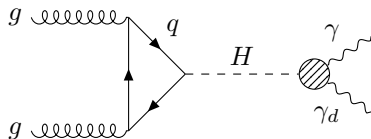
- 1 well-identified ("tight") isolated photon with  $p_T^\gamma > 50$  GeV;





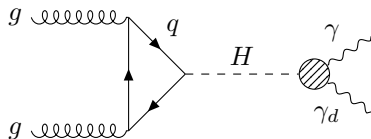
**Signal:**  $gg \rightarrow H \rightarrow \gamma\gamma_d$

- 1 well-identified ("tight") isolated photon with  $p_T^\gamma > 50$  GeV;
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- 1 well-identified ("tight") isolated photon with  $p_T^\gamma > 50$  GeV;
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- leptons veto,  $N_l = 0, l \in \{e, \mu, \tau\}$ ;



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- missing transverse momentum  $p_T^{\text{miss}} > 100$  GeV;

- leptons veto,  $N_l = 0, l \in \{e, \mu, \tau\}$ ;
- transverse mass  $m_T > 80$  GeV.

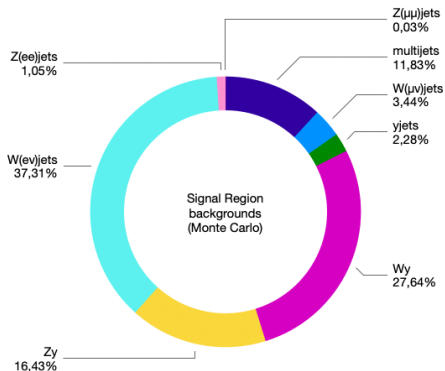
$$m_T = \sqrt{2p_T^{\text{miss}} p_T^\gamma (1 - \cos \Delta\Phi(\vec{p}_T^\gamma, \vec{p}_T^{\text{miss}}))}$$

## Background processes:

- irreducible:  $Z(\rightarrow \nu\nu)\gamma$ ;
- reducible:
  - $W(\rightarrow l\nu_l)\gamma$ , lost lepton;
  - multijets,  $W(\rightarrow \mu\nu_\mu)\text{jets}$ ,  $Z(\rightarrow \mu\mu)\text{jets}$ , jets faking photons ( $\text{jet} \rightsquigarrow \gamma$ );
  - $\gamma\text{jets}$ , true photon but fake  $p_T^{\text{miss}}$ ;
  - $W(\rightarrow e\nu_e)\text{jets}$ ,  $Z(\rightarrow ee)\text{jets}$ , electrons faking photons ( $e \rightsquigarrow \gamma$ ).

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  - $\gamma$ jets, true photon but fake  $p_T^{\text{miss}}$ ;
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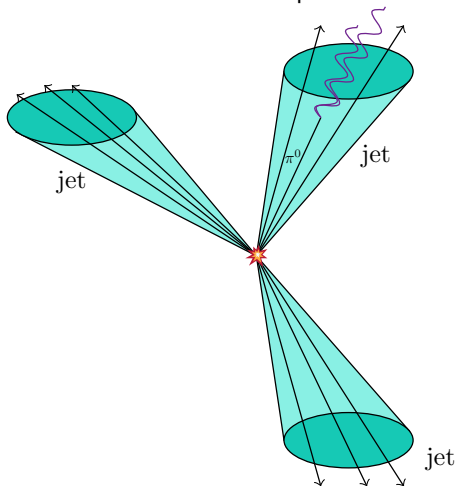
**Background  
composition in pure  
Monte Carlo samples.**



# Jets faking photons background

# Jets faking photons

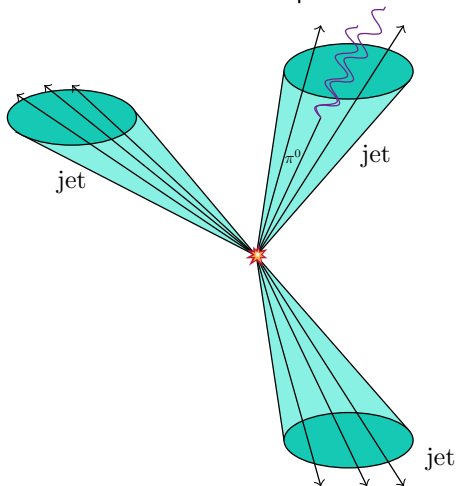
Before the reconstruction process:



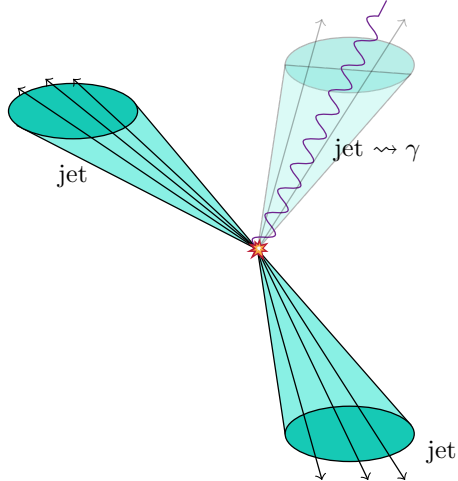


# Jets faking photons

Before the reconstruction process:



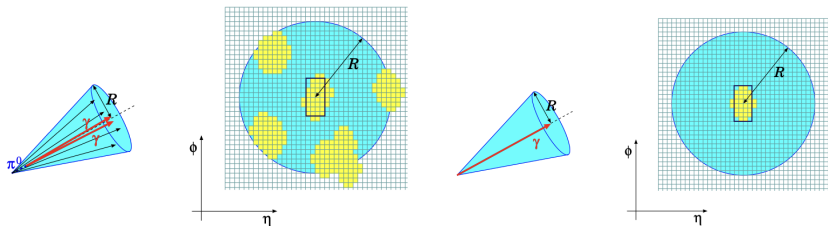
After the reconstruction process:



How can we discriminate true and fake photons?

⇒ **Isolation**

$$isol = \frac{E_T^{isolation}}{p_T^\gamma}$$

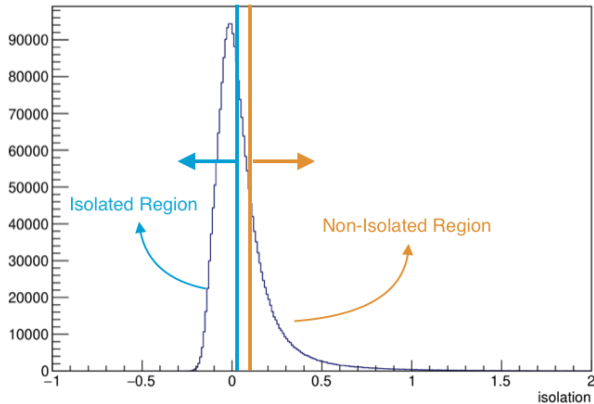


# Isolation Regions

How can we discriminate true and fake photons?

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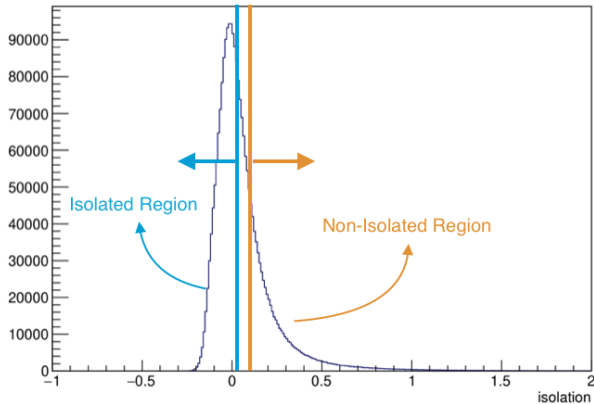


# Fake factors

How can we discriminate true and fake photons?

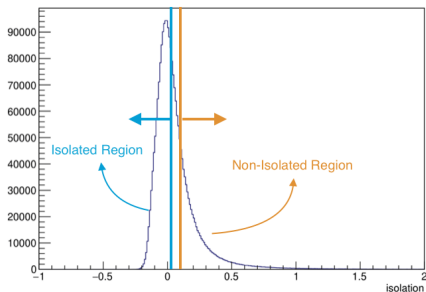
⇒ **Isolation**

$$isol = \frac{E_T^{isolation}}{p_T^\gamma}$$



$$f = \left( \frac{N_{j \rightarrow \gamma}^{isol}}{N_{j \rightarrow \gamma}^{non-isol}} \right)_{tight}$$

# Fake factors calculation



$$f = \left( \frac{N_{j \rightarrow \gamma}^{isol}}{N_{j \rightarrow \gamma}^{non-isol}} \right)_{tight}$$

## Problems:

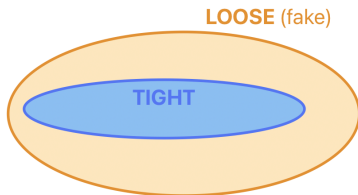
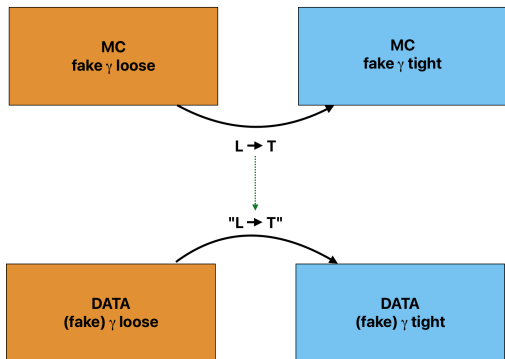
- We cannot fully trust **Monte Carlo** as isolation is typically not well modelled for jets faking photons;
- In **data**, we cannot distinguish true and fake photons!

⇒ we need to **extract** the tight fake photons isolation distribution in data.

# Extrapolation method

How to get tight fake photons isolation distribution in data?

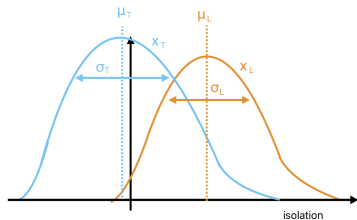
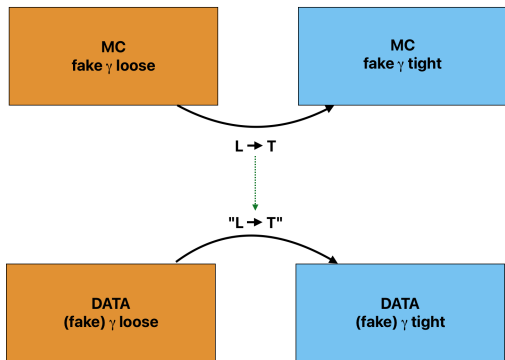
**L: loose**  
**T: tight**



# Extrapolation method

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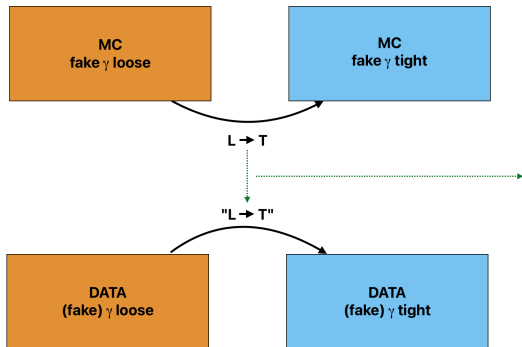
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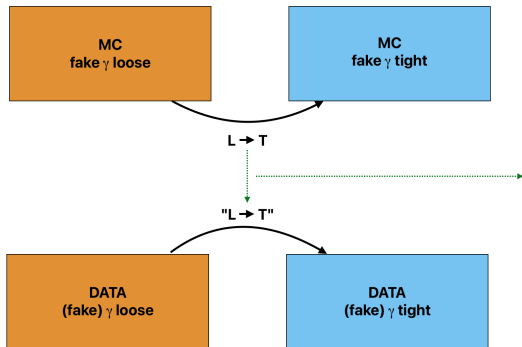
$$\frac{\sigma_T^{data}}{\sigma_L^{data}} = \frac{\sigma_T^{MC}}{\sigma_L^{MC}}$$
$$\frac{\mu_T^{data} - \mu_L^{data}}{\sigma_L^{data}} = \frac{\mu_T^{MC} - \mu_L^{MC}}{\sigma_L^{MC}}$$



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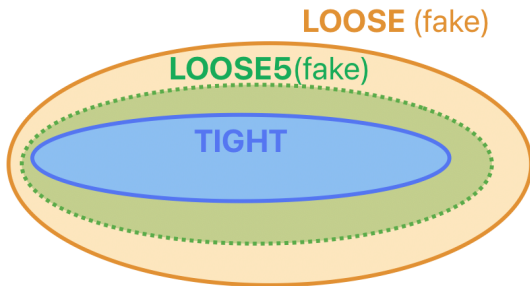


$$\frac{\sigma_T^{data}}{\sigma_L^{data}} = \frac{\sigma_T^{MC}}{\sigma_L^{MC}}$$
$$\frac{\mu_T^{data} - \mu_L^{data}}{\sigma_L^{data}} = \frac{\mu_T^{MC} - \mu_L^{MC}}{\sigma_L^{MC}}$$

$\rightarrow$  to be validated

# Validation of the hypotheses

**L: loose**  
**T: tight**  
**L5: loose5**

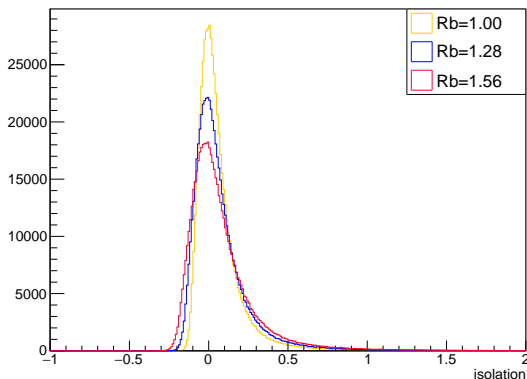


$$\frac{\sigma_{L5}^{data}}{\sigma_L^{data}} = R_b \frac{\sigma_{L5}^{MC}}{\sigma_L^{MC}}$$
$$\frac{\mu_{L5}^{data} - \mu_L^{data}}{\sigma_L^{data}} = R \frac{\mu_{L5}^{MC} - \mu_L^{MC}}{\sigma_L^{MC}}$$

→ validated on another photon sample similar to loose (loose5)

# Extrapolated distributions

Extrapolated isolation distribution of fake photons passing tight identification

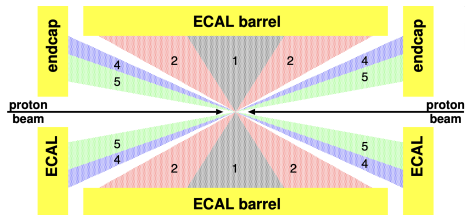


Uncertainties on  $R$ ,  $R_b$  have been assumed to be equal to  $|R - 1|$ ,  $|R_b - 1|$ .

# Fake factors

Fake factors have been computed in different **geometric regions** of the detector...

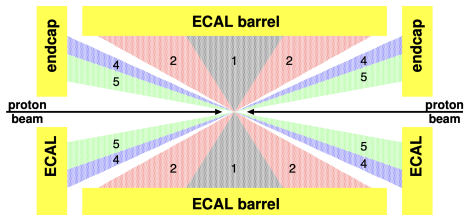
Region	f	$\sigma_f^{sys}$	%
1	1.51	0.18	11.7 %
2	2.03	0.35	17.1 %
4	1.95	0.34	17.2 %
5	1.70	0.27	15.7 %



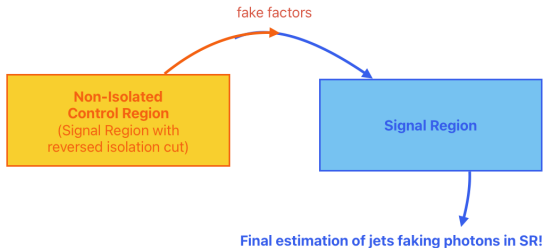
# Jets faking photons final estimation

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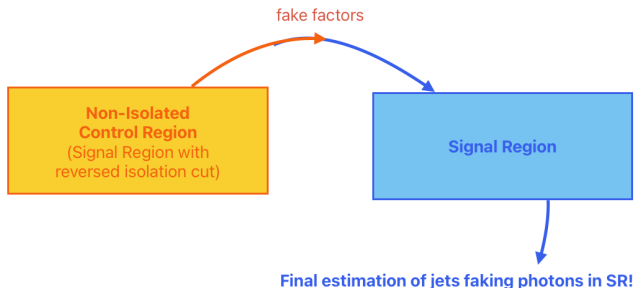


...and then **applied** to the Non-Isolated Control Region.



# Jets faking photons final estimation

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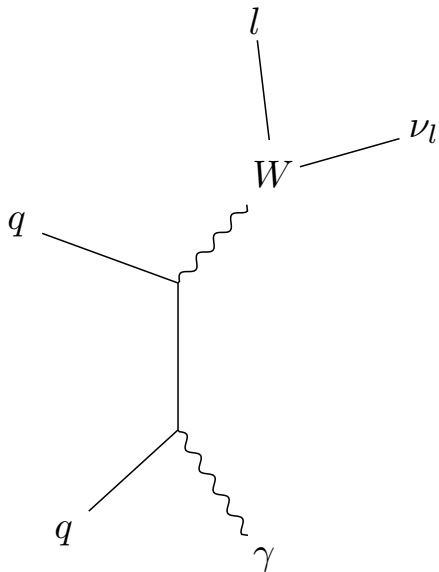
Comparison with the yield in Monte Carlo:

$$N_{j \rightsquigarrow \gamma}^{MC} = 228 \pm 70$$

$$N_{j \rightsquigarrow \gamma}^{data-driven} = 718 \pm 116$$

# $W\gamma$ background

# $W\gamma$ background



$W(\rightarrow l\nu_l)\gamma$

- 1 photon
- genuine  $p_T^{miss}$

→ when the lepton gets lost, this is a background for our analysis



We need to extract **K-factors** to be used to correct the cross-section approximation of Monte Carlo simulations.

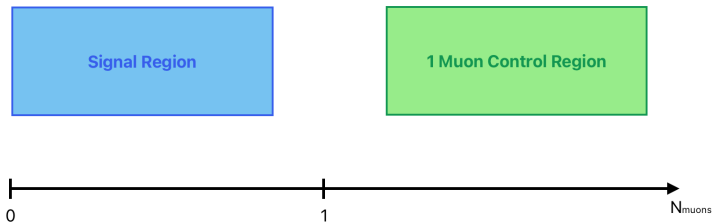
$$K = \left( \frac{N_{W\gamma}^{data}}{N_{W\gamma}^{MC}} \right)_{1\mu CR}$$

# K-factors and 1 Muon Control Region

We need to extract **K-factors** to be used to correct the cross-section approximation of Monte Carlo simulations.

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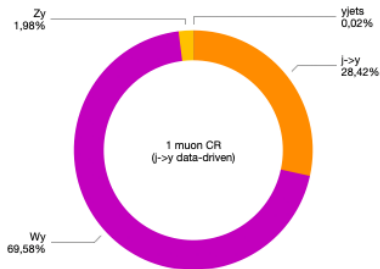
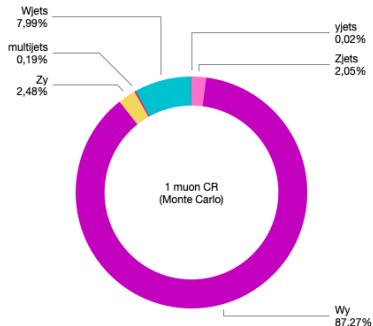
We construct a **1 Muon Control Region** enriched of  $W\gamma$  events, where the muon is treated as **invisible**.



# Jets faking photons in the 1 Muon Control Region

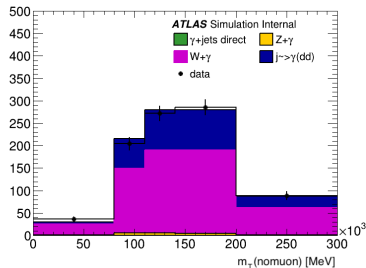
The jets faking photons contribution to the 1 Muon Control Region is estimated using the **fake factors** calculated in the previous part of the work.

→ Jets faking photons data-driven estimation is much higher than Monte Carlo!



# K-factors calculation

K-factors have been computed in different **transverse mass bins**...

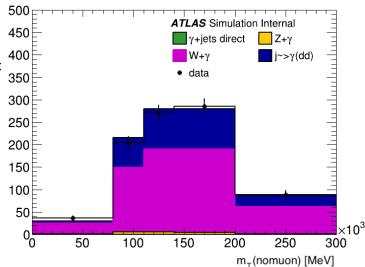


$m_T$  without muon contribution  
in the 1 Muon CR for  $W\gamma$ ,  $Z\gamma$ ,  
 $\gamma$ jets and data.

# K-factors calculation

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$$K = \left( \frac{N_{W\gamma}^{data}}{N_{W\gamma}^{MC}} \right)_{1\mu CR} = \left( \frac{N^{data} - N_{bkg \neq W\gamma}}{N_{W\gamma}^{MC}} \right)_{1\mu CF}$$

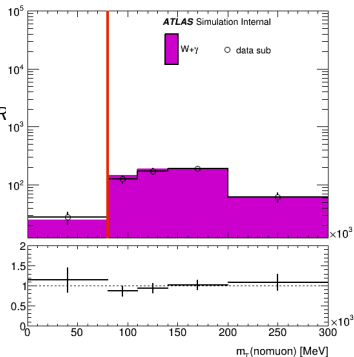


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$m_T$  without muon contribution  
in the 1 Muon CR for  $W\gamma$  and  
subtracted data.

# K-factors calculation

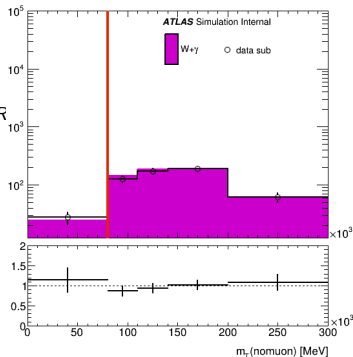
K-factors have been computed in different **transverse mass bins**...

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$m_T$ (GeV)	K	$\sigma_K^{stat}$	$\sigma_K^{sys}$	$\sigma_K^{tot}$
80-110	0.869	0.131	0.074	0.150
110-140	0.939	0.119	0.076	0.141
140-200	1.023	0.117	0.073	0.138
>200	1.089	0.197	0.067	0.208

where

$$\sigma_K^{sys} = \frac{K(f - \sigma_f) - K(f + \sigma_f)}{2}$$



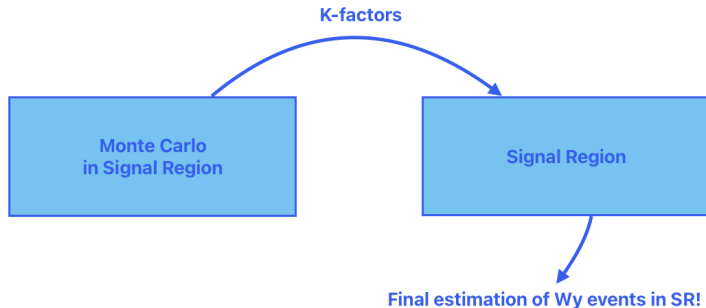
$m_T$  without muon contribution in the 1 Muon CR for  $W\gamma$  and subtracted data.

# K-factors application

K-factors have been computed in different **transverse mass** bins...

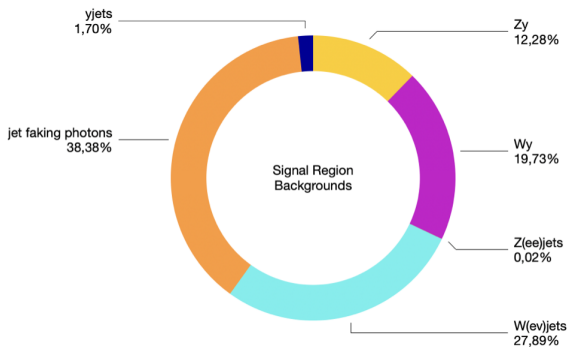
$m_T$ (GeV)	K	$\sigma_K^{tot}$	%
80-110	0.869	0.150	17.3 %
110-140	0.939	0.141	15.0 %
140-200	1.023	0.138	13.5 %
>200	1.089	0.208	19.1 %

..and then **applied** to Monte Carlo in Signal Region.

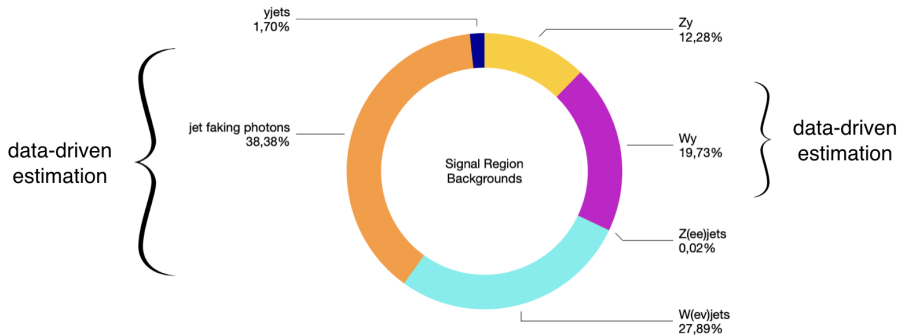




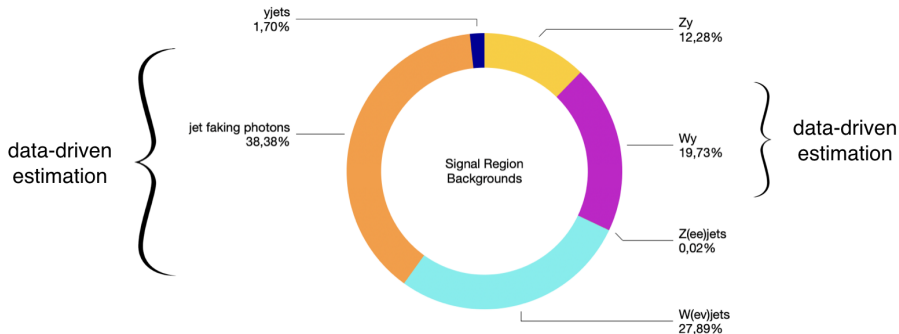
# Backgrounds in Signal Region



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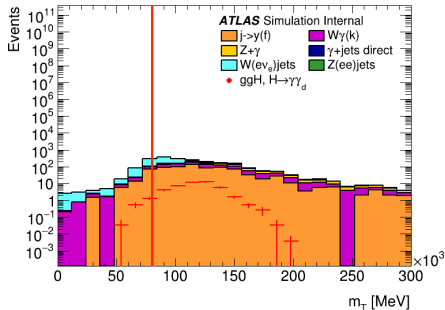
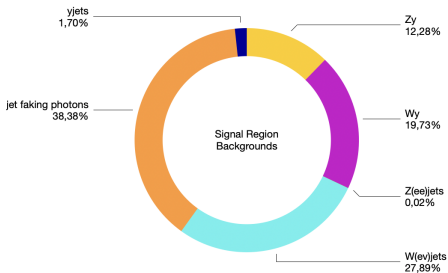


# Backgrounds in Signal Region



Jets faking photons and  $W\gamma$  constitute  $\sim 60\%$  of the total background in Signal Region.

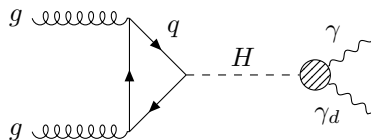
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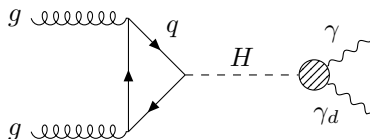
# Conclusions

- I contributed to the backgrounds estimation in the ATLAS search for dark photons;



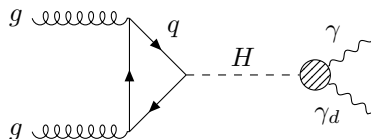
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- I contributed to the backgrounds estimation in the ATLAS search for dark photons;
- I developed a new data-driven technique for the jets faking photons estimation;



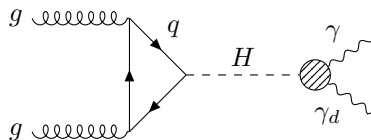
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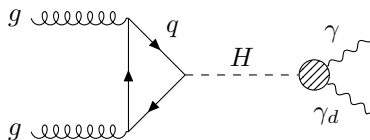
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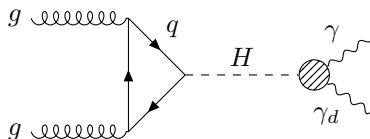
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- I developed a new data-driven technique for the jets faking photons estimation;
- I estimated the  $W\gamma$  background, using the k-factors method;
- Jets faking photons and  $W\gamma$  constitute  $\sim 60\%$  of the total background in Signal Region;
- The analysis  $gg \rightarrow H \rightarrow \gamma\gamma_d$  is on-going;



# Conclusions

- I contributed to the backgrounds estimation in the ATLAS search for dark photons;
- I developed a new data-driven technique for the jets faking photons estimation;
- I estimated the  $W\gamma$  background, using the k-factors method;
- Jets faking photons and  $W\gamma$  constitute  $\sim 60\%$  of the total background in Signal Region;
- The analysis  $gg \rightarrow H \rightarrow \gamma\gamma_d$  is on-going;
- These results will enter the official ATLAS analysis publication.



# Backup

# Dark Matter, Dark Sector, Dark Photon

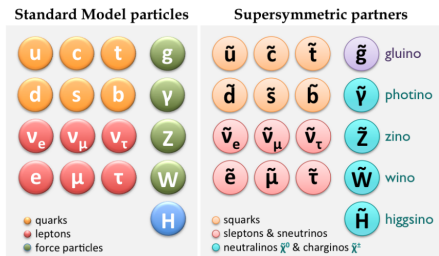
DM candidates should be:

- neutral;
- cold, non-relativistic at the time of CMB formation;
- stable or at least with lifetime longer than the age of the Universe;
- weakly interacting with themselves and with ordinary matter.

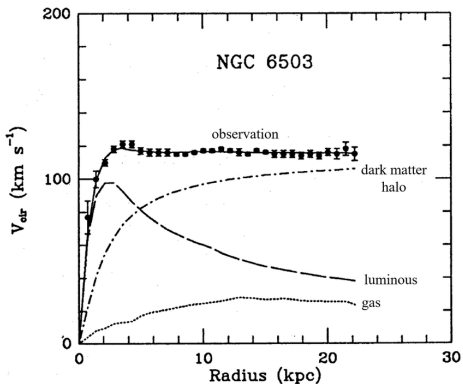
Some candidates:

- WIMPs, Weakly Interacting Massive Particles, e.g. SUSY;
- sterile neutrinos, RH neutrinos with low mixing constant with ordinary neutrinos;
- many others...

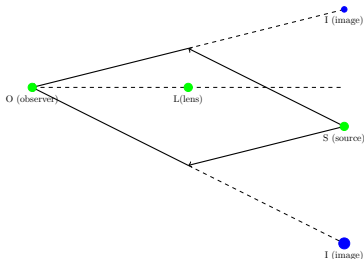
- SUSY theory introduced to explain the difference between the measured value of the Higgs boson mass and the one predicted by the first order calculation, including the top annihilation term;
- particles with spin differing by half a unit with respect to the SM;
- s-top annihilation term would compensate the top annihilation term;
- viable DM candidates: neutralinos.



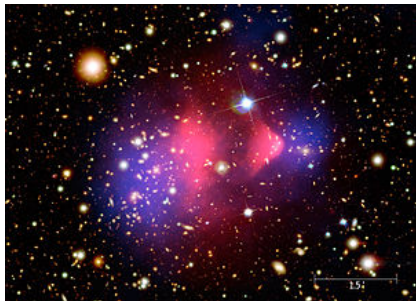
- Rotational curves of galaxies



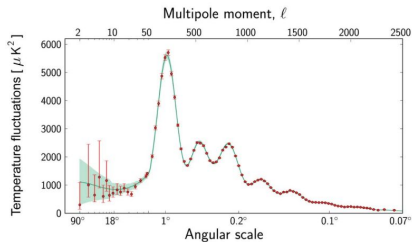
- Gravitational lensing



- Bullet cluster



- CMB spectrum





# A portal to the dark sector

The portal and can take various forms:

- vector portal  $\implies$  massive dark photon

$$\mathcal{L}_{kin.mix.} = \frac{1}{2} \varepsilon F_{\mu\nu} F'^{\mu\nu}$$

where  $F, F'$  are field strength tensors of the SM  $U(1)$  and the dark  $U(1)_D$ . For a massless dark photon, the direct kinetic mixing is not possible. There should be "something" in between.

- scalar portal;
- neutrino portal.

# Analysis

Data used in this work of thesis are normalized at:

$$\mathcal{L} = 25,76 \text{ fb}^{-1}$$

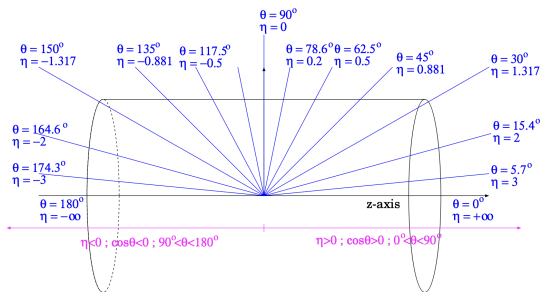
Normalizing data with full Run3 luminosity (expected  $\sim 300 \text{ fb}^{-1}$ ) will increase the statistic and reduce statistical uncertainties.

# Pseudorapidity

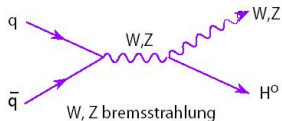
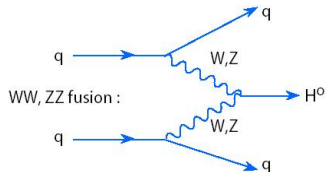
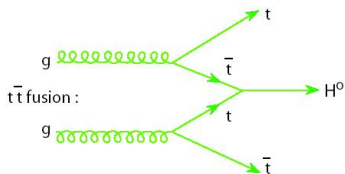
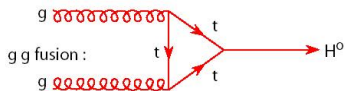
$$\eta = -\ln \left[ \tan \frac{\theta}{2} \right]$$

where  $\theta$  is the polar angle. This formula set a one-to-one correspondence between the  $\theta$  coordinate of a polar system and  $\eta$ , moving domain from  $(0, \pi)$  to  $(-\infty, +\infty)$ .

$\Delta\eta$  is invariant under Lorentzian boosts along beam axis; this becomes important as the reference frame of the center-of-mass of the interaction is unknown.



# Higgs production channels



## All the cuts defining Signal Region

- $n_e = 0$ , electrons veto;
- $n_\mu = 0$ , muon veto;
- $n_\tau = 0$ , tau lepton veto;
- $p_T^{miss} > 100$  GeV;
- $n_\gamma^{isol} = 1$ , one isolated photon;
- $p_T^\gamma > 50$  GeV;
- $m_T > 80$  GeV;
- $n_{jet} \leq 3$ , maximum 3 jets;
- $\Delta\Phi(\vec{p}_T^{miss}, [\vec{p}_T^{miss}]_\gamma) \geq 1.25$ ,  $\gamma, \gamma_d$  in the transverse region;
- $S_{p_T^{miss}} > 6$ , in order to remove fake  $p_T^{miss}$ ;
- $\Delta p_T^{miss} > -10$  GeV
- $|\eta_\gamma| < 1.75$ ,  $\gamma, \gamma_d$  in the transverse region;;
- $\Delta\Phi(\vec{p}_T^{miss}, [\vec{p}_T^{miss}]_j) \leq 0.75$ , the Higgs boson should scatter on the jets;
- $\Delta\Phi(\vec{p}_T^{j1}, \vec{p}_T^{j2}) \leq 2.5$ , in order to remove dijets.

$$\Delta |\vec{p}_T^{miss}| = |[\vec{p}_T^{miss}]_{noJVT}| - |\vec{p}_T^{miss}|$$

This cut targets  $\gamma$ +jets background events.

As the **hard scattering event vertex** is chosen by picking the one with the highest scalar sum of the momenta of all the tracks produced in it, in a  $\gamma$ +jets event, where a big portion of momentum is carried away by the photon, there is a non-negligible probability to elect a pile-up vertex as hard scattering vertex.

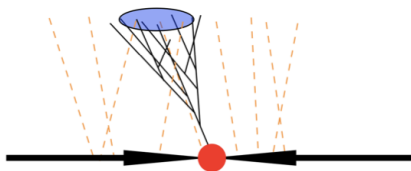
If JVT cut is applied in such a case, this will lead to **exclude the real jet** from  $\vec{p}_T^{miss}$  calculation, hence resulting in a large fake missing transverse momentum in the final state.

Events where the  $p_T^{miss}$  calculated with JVT is much higher than the  $p_T^{miss}$  calculated with JVT are in most of the cases events with a mis-reconstructed primary vertex and can be then **excluded**.

## Jet Vertex Tag (JVT)

$$JVT = \frac{\sum_{j \in \text{hard scattering}} p_T^j}{\sum_j p_T^j}$$

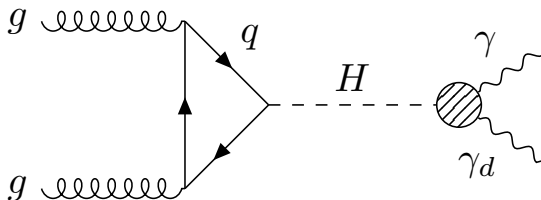
⇒ reject the jet if JVT is under a certain threshold!





# Analysis Trigger

- 1 tight photon,  $N_\gamma = 1$ ;
- photon transverse momentum  $|\vec{p}_T^\gamma| > 50$  GeV;
- missing transverse momentum  $|\vec{p}_T^{miss}| > 40$  GeV with calculation based on cells and  $|\vec{p}_T^{miss}| > 70$  GeV with calculation including tracks;
- transverse mass  $m_T > 80$  GeV.



# Backgrounds estimation strategy

- irreducible background:  $Z(\rightarrow \nu\nu)\gamma$ ;
- lost lepton:  $W(\rightarrow l\nu_l)\gamma$ ;
- jets faking photons: multijets, Zjets, Wjets;
- electrons faking photons:  $W(\rightarrow e\nu_e)$ jets,  $Z(\rightarrow ee)$ jets;
- fake  $p_T^{miss}$ : multijets,  $\gamma$ jets.

## Strategy 1

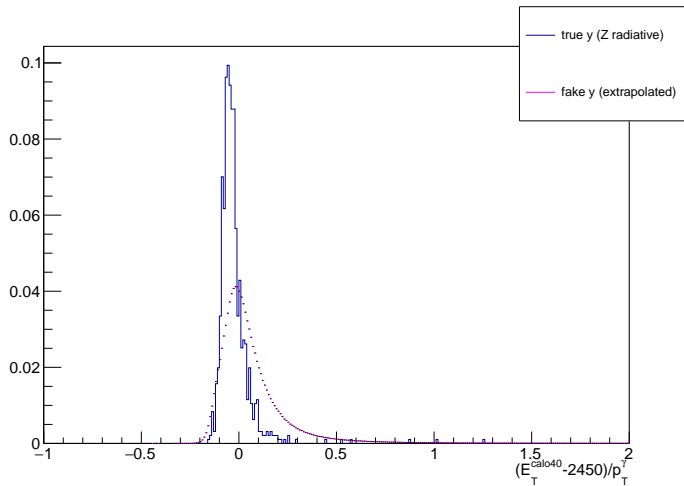
- electrons faking photons data-driven
- jets faking photons data-driven
- $W\gamma, Z\gamma$  from leptons CR
- $\gamma$ jets Monte Carlo

## Strategy 2

- electrons faking photons data-driven
- fake  $p_T^{miss}$  data-driven
- $W\gamma + W$ jets,  $Z\gamma + Z$ jets from leptons CR

# Jets faking photons

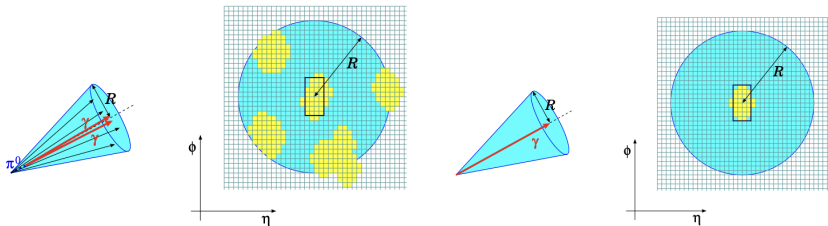
# True and fake photons isolation comparison



# Calorimeter relative isolation

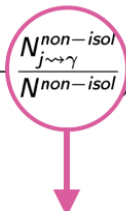
$$isol_{calo}^{rel} = \frac{E_T^{calo40} - 2450}{p_T^\gamma}$$

where  
 $p_T^\gamma$  is the photon transverse momentum;  
 $E_T^{calo40}$  is the energy not belonging to the photon measured in a cone with radius  $\Delta R = 0.4$  around the photon;  
2,45 GeV is a pedestal factor.



The fraction of true photons in the Non-Isolated Region is given by the **purity**  $P$ .

$$P = \left( \frac{N_{\gamma}^{non-isol}}{N^{non-isol}} \right)_{tight} = \left( \frac{N^{non-isol} - N_{j \rightsquigarrow \gamma}^{non-isol}}{N^{non-isol}} \right)_{tight} = \left( 1 - \frac{N_{j \rightsquigarrow \gamma}^{non-isol}}{N^{non-isol}} \right)_{tight}$$



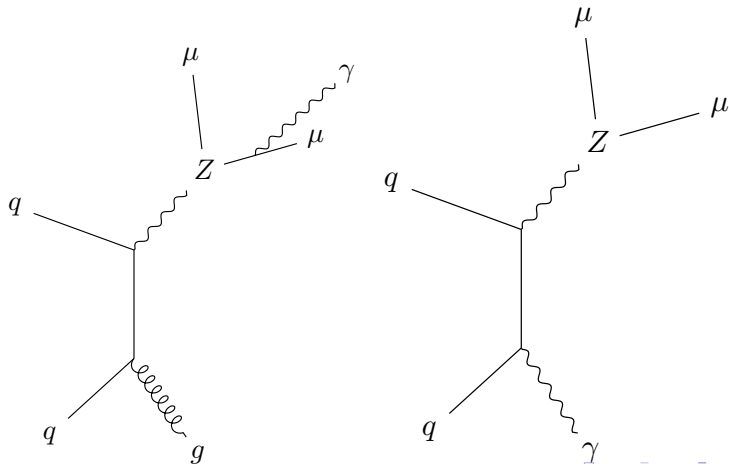
depends on the number of non isolated photons produced in the event.

It is not an intrinsic property of how we "see" jets!  
We would like to have  $P \sim 0$ , i.e. no contamination of true photons in the Non-Isolated Region.

# Non-Isolated Region definition

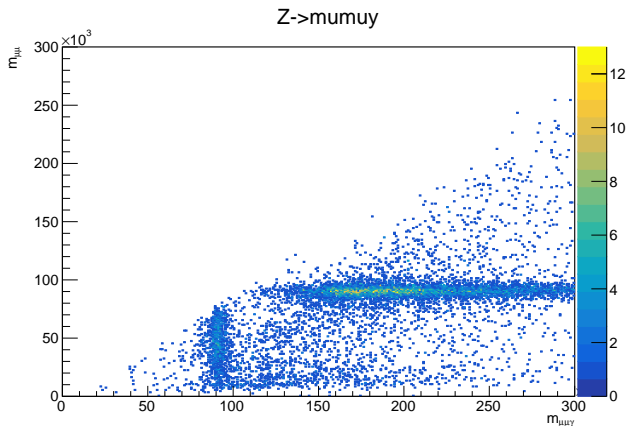
Let's define the Non-Isolated Region such to have  $P \sim 0$ , i.e. no contamination of true photons in the Non-Isolated Region.

Let's look at a **pure** sample of photons, that can be obtained selecting  $Z(\rightarrow \mu\mu\gamma)$  events.



# Non-Isolated Region definition

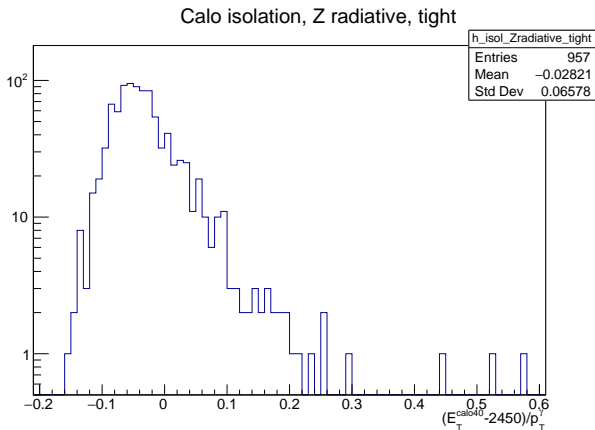
In order to select  $Z(\rightarrow \mu\mu\gamma)$  events, let's consider events in the  $\mu\mu\gamma$  sample with a tight photon and  $80 \text{ GeV} < m_{\mu\mu\gamma} < 100 \text{ GeV}$ .





# Non-Isolated Region definition

Looking at the calorimeter relative isolation of these events, we decide to put the cut of the Non-Isolated Region at 0.1.



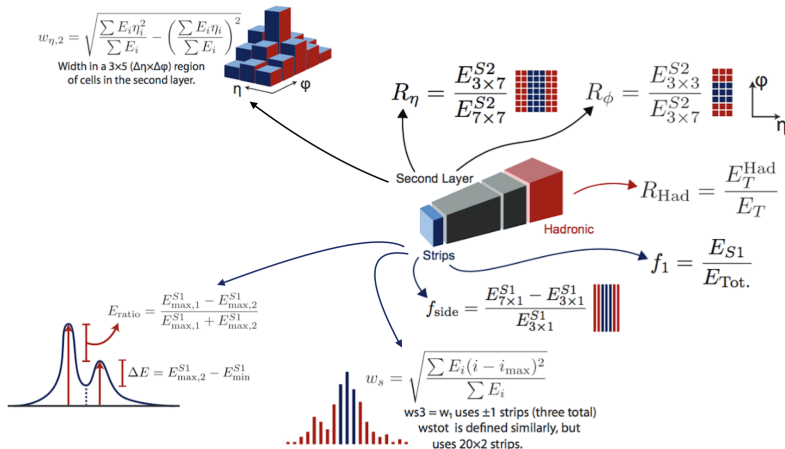
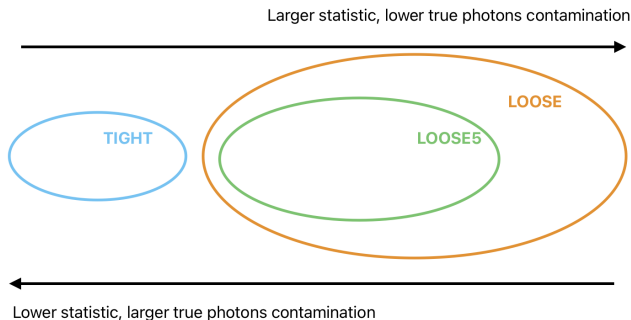


Figure 3: Discriminant Variables (DVs) describing shower shapes, energy ratios and width of the energy deposit

Different possible **ID selection**:

- **tight**, passing tight cuts on all the DVs;
- **loose**, if they pass looser cuts on some DVs ( $R_\eta$ ,  $R_{had}$  and  $w_{\eta,2}$ ) but not the tight ones;
- **loose5**, if they are loose and pass tight cuts on more DVs ( $R_\eta$ ,  $R_{had}$ ,  $w_{\eta,2}$  and  $R_\phi$ );



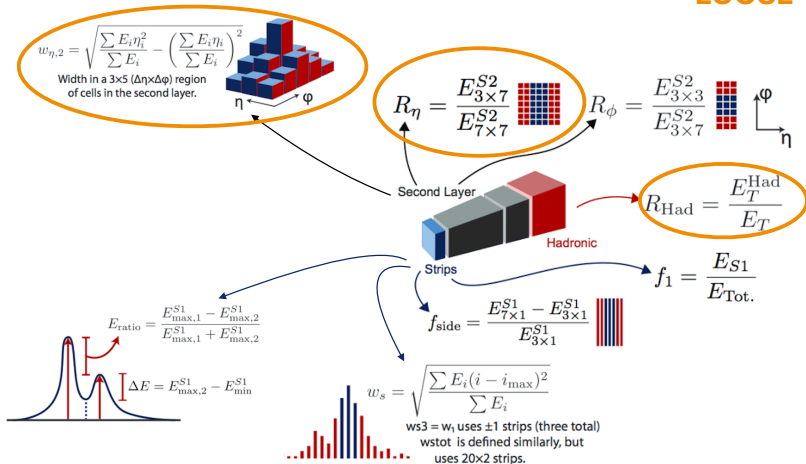


Figure 4: Discriminant Variables (DVs) describing shower shapes, energy ratios and width of the energy deposit (loose)

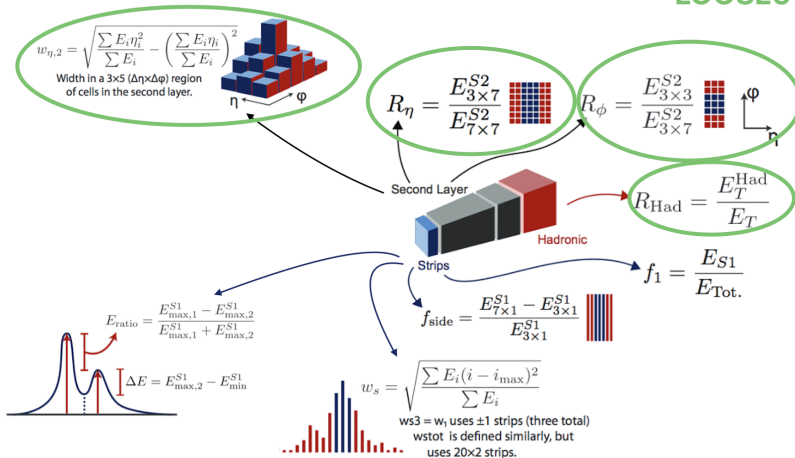


Figure 5: Discriminant Variables (DVs) describing shower shapes, energy ratios and width of the energy deposit (loose5)

# Step 1: get L->T transformation from MC

Let's assume that tight  $isol_T^{MC}$  and loose  $isol_L^{MC}$  distributions in MC are linked by an affine transformation.

$\mu$ : **median**  
 $\sigma$ : **width**

$$isol_T^{MC} = a + b isol_L^{MC}$$

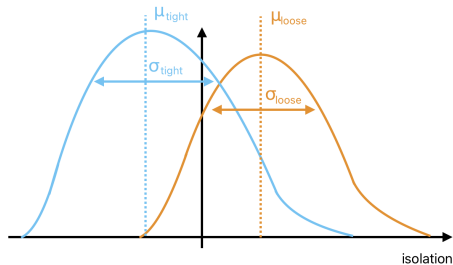
We want to find  $a, b$  such that:

$$isol_T^{MC} = \mu_T^{MC} + \frac{\sigma_T^{MC}}{\sigma_L^{MC}} (isol_L^{MC} - \mu_L^{MC})$$

so

$$a = \mu_T^{MC} - \frac{\sigma_T^{MC}}{\sigma_L^{MC}} \mu_L^{MC}$$

$$b = \frac{\sigma_T^{MC}}{\sigma_L^{MC}}$$



# Step 1: get L->T transformation from MC

Let's assume:

$\mu$ : median

$\sigma$ : width

- the scale factor  $b$  stays the same in MC and data;

$$\frac{\sigma_T^{data}}{\sigma_L^{data}} = \frac{\sigma_T^{MC}}{\sigma_L^{MC}}$$

- the offset  $a$  in data should depend on  $\sigma_L^{data}$ ,  $\sigma_T^{data}$ , which is known, and on  $\mu_T^{data}$ , which is unknown. So we assume the shift of the average going from loose to tight is proportional to the rms in both data and MC.

$$a = \mu_T^{MC} - \frac{\sigma_T^{MC}}{\sigma_L^{MC}} \mu_L^{MC} \quad \longrightarrow \quad a = \mu_T^{DATA} - \frac{\sigma_T^{DATA}}{\sigma_L^{DATA}} \mu_L^{DATA}$$

known

$$\frac{\mu_T^{data} - \mu_L^{data}}{\sigma_L^{data}} = \frac{\mu_T^{MC} - \mu_L^{MC}}{\sigma_L^{MC}}$$
$$\mu_T^{data} = \mu_L^{data} + \frac{\sigma_L^{data}}{\sigma_L^{MC}} (\mu_T^{MC} - \mu_L^{MC})$$

## Step 2: apply L->T transformation to DATA

$\mu$ : **median**  
 $\sigma$ : **width**

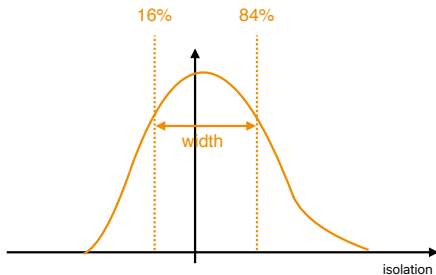
Putting all together, the transformation for DATA is:

$$isol_T^{data} = \mu_L^{data} + \frac{\sigma_L^{data}}{\sigma_L^{MC}} (\mu_T^{MC} - \mu_L^{MC}) + \frac{\sigma_T^{MC}}{\sigma_L^{MC}} (isol_L^{data} - \mu_L^{data})$$

→ We obtain **tight fake photons** distributions in DATA.



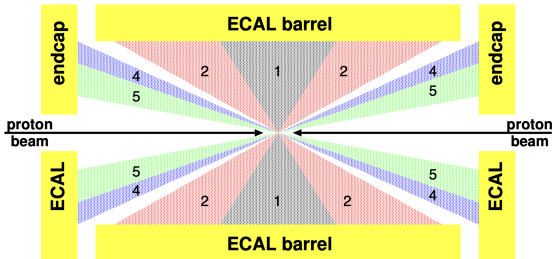
Width is calculated as difference between the 16th and 84th percentile.



# Binning in $\eta$

Pseudorapidity binning is chosen considering the detector geometry:

- Region 1:  $[0; 0.6]$ , the upper limit  $\eta = 0.6$  is the point after which the material in front of ECAL increases a lot;
- Region 2:  $[0.6; 1.37]$ , the upper limit is defined by the beginning of the crack region;
- Crack region:  $[1.37; 1.52]$ , not used due to low reconstruction performance;
- Region 4:  $[1.52; 1.81]$ , the upper limit is the point where the presampler ends;
- Region 5:  $[1.81; 2.37]$ .



# Analysis Trigger, $R$ , $R_b$

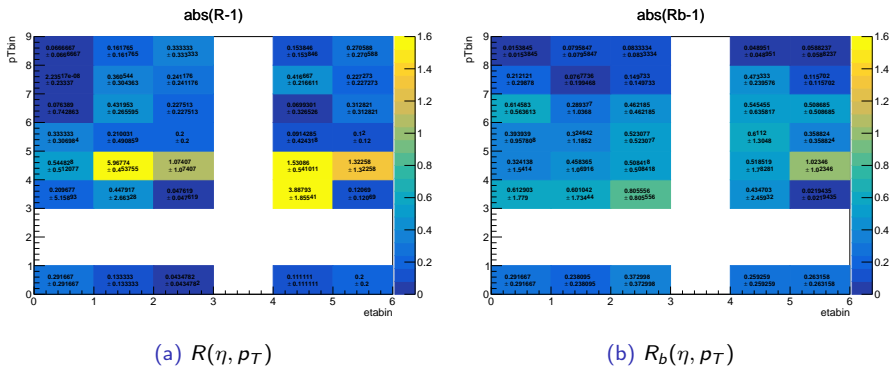


Figure 6:  $R(\eta, p_T)$  and  $R_b(\eta, p_T)$  computed with  $(\mu_{\text{med}}, \sigma_{q16})$  using Analysis Trigger.

# Median and width

We decide to use (median&width) instead of (mean&sigma) for the extrapolation:  
 $R, R_b$  values in different  $\eta, p_T$  are indeed less spread.

Trigger	Ratio	Used variables	RMS in $\eta, p_T$ bins
Analysis	$R$	$\mu, \sigma$	0.19
		$\mu_{med}, \sigma_{q16}$	0.14
	$R_b$	$\mu, \sigma$	0.30
		$\mu_{med}, \sigma_{q16}$	0.28
$p_T^{miss}$	$R$	$\mu, \sigma$	0.27
		$\mu_{med}, \sigma_{q16}$	0.10
	$R_b$	$\mu, \sigma$	0.22
		$\mu_{med}, \sigma_{q16}$	0.13
Leptonic	$R$	$\mu, \sigma$	0.54
		$\mu_{med}, \sigma_{q16}$	0.18
	$R_b$	$\mu, \sigma$	0.51
		$\mu_{med}, \sigma_{q16}$	0.36

# Inclusive Region and Trigger

**Problem:**  $R$  and  $R_b$  are quite unstable. It is not possible to perform the extrapolation in exclusive regions in  $p_T, \eta$ .

**Solution:** Let's be either inclusive in  $p_T$  or in  $\eta$ .

# Inclusive Region and Trigger

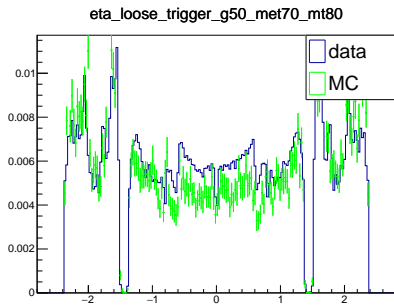
**Problem:** How to choose the Inclusive Region and the Trigger to be used?

Trigger ↓ Inclusivity	Analysis	$p_T^{miss}$	Leptonic
$p_T$			
$\eta$			

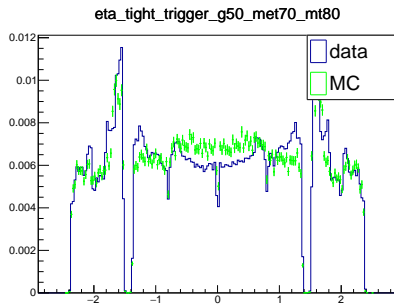
**Solution:** Let's choose the configuration satisfying the requirements:

- the spectrum of the inclusive variable in data and MC should be similar;
- the configuration should ensure the lowest uncertainties on the fake factors.

# $\eta, p_T$ spectra in data and MC



(a) Loose



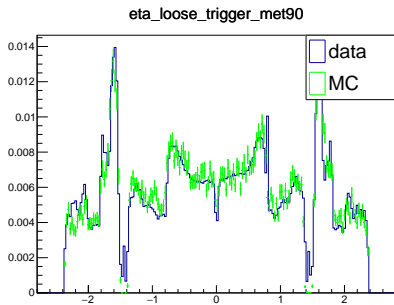
(b) Tight

Figure 7:  $\eta$  distribution for loose and tight  $\gamma$  in data and MC samples (Analysis Trigger).

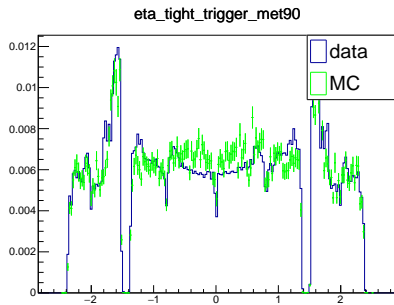
Trigger ↓ Analysis	Analysis	$p_T^{mix}$	Leptonic
$p_T$			
$\eta$			

⇒ good agreement ✓

# $\eta$ , $p_T$ spectra in data and MC



(a) Loose



(b) Tight

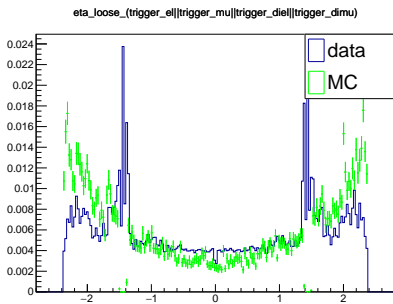
Figure 8:  $\eta$  distribution for loose and tight  $\gamma$  in data and MC samples ( $p_T^{miss}$  Trigger).

Trigger Inclusion	Analysis	$p_T^{miss}$	Leptonic
$p_T$			
$\eta$			

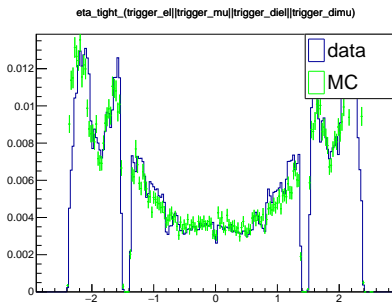
$\Rightarrow$  good agreement ✓



# $\eta$ , $p_T$ spectra in data and MC



(a) Loose



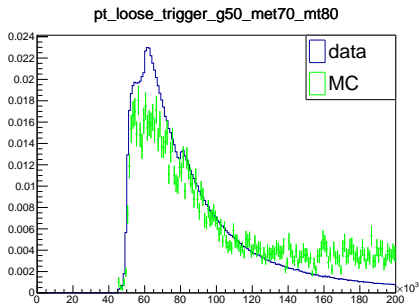
(b) Tight

Figure 9:  $\eta$  distribution for loose and tight  $\gamma$  in data and MC samples (Leptonic Trigger).

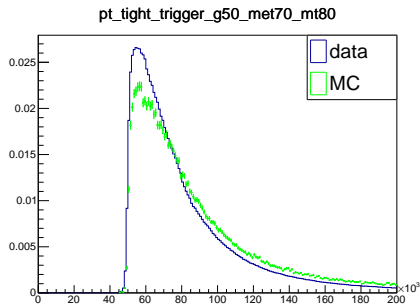
trigger ↓ recovery	Analysis	$p_T^{\text{miss}}$	Leptonic
$p_T$			
$\eta$			

⇒ bad agreement ✘

# $\eta, p_T$ spectra in data and MC



(a) Loose



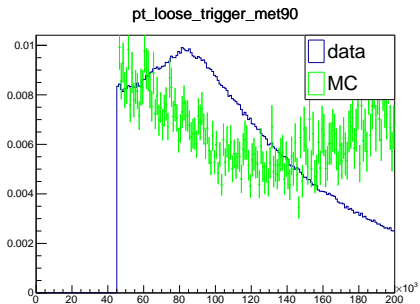
(b) Tight

Figure 10:  $p_T$  distribution for loose and tight  $\gamma$  in data and MC (Analysis Trigger).

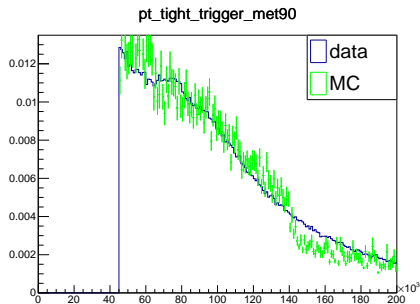
Trigger → Technique	Analysis	$p_T^{\text{miss}}$	Leptonic
$p_T$			
$\eta$			

⇒ good agreement ✓

# $\eta$ , $p_T$ spectra in data and MC



(a) Loose



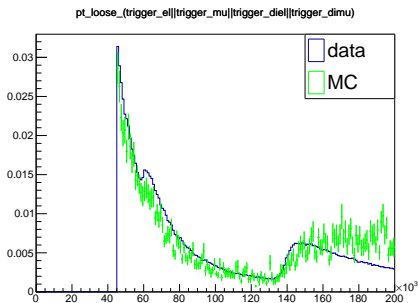
(b) Tight

Figure 11:  $p_T$  distribution for loose and tight  $\gamma$  in data and MC ( $p_T^{miss}$  Trigger).

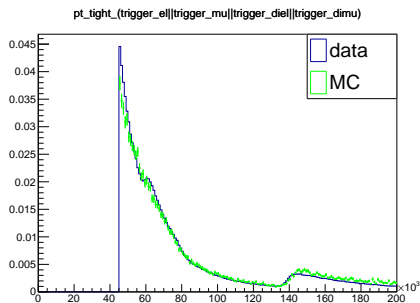
Trigger Analysis	$p_T^{miss}$	Leptonic
$p_T$		
$\eta$		

$\Rightarrow$  bad agreement ✗

# $\eta, p_T$ spectra in data and MC



(a) Loose



(b) Tight

Figure 12:  $p_T$  distribution for loose and tight  $\gamma$  in data and MC (Leptonic Trigger).

trigger ↓ technique	Analysis	$p_T^{\text{miss}}$	Leptonic
$p_T$			
$\eta$			

$\Rightarrow$  good agreement ✓

# Inclusive Region and Trigger

Trigger → ↓ Inclusivity	Analysis	$p_T^{miss}$	Leptonic
$p_T$			
$\eta$			

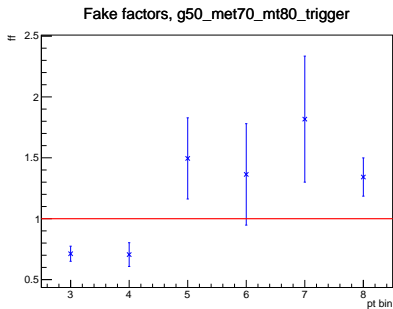
# $R, R_b$ in different configurations

Trigger	Incl. Region	$R_{nom}$	$R_{up}$	$R_{down}$	$R_{b,nom}$	$R_{b,up}$	$R_{b,down}$
analysis	$\eta$	1.10	1.20	1.00	1.36	1.72	1.00
	$p_T$	1.06	1.12	1.00	1.28	1.56	1.00
$p_T^{miss}$	$\eta$	0.82	1.00	0.64	1.06	1.12	1.00
leptonic	$p_T$	0.79	1.00	0.58	1.02	1.04	1.00

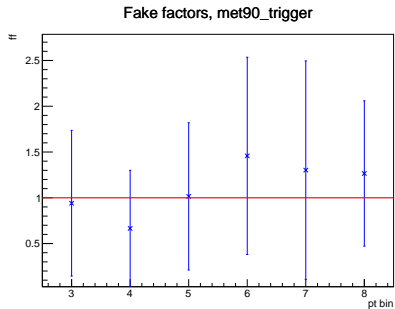
Table 1: Nominal and varied values of  $R$  and  $R_b$  for the different configurations of inclusive region and trigger possible.

Let's now extrapolate the tight fake photons isolation distributions in data using these  $R, R_b$  (nominal and varied) for each configuration.

# Fake factors in different configurations



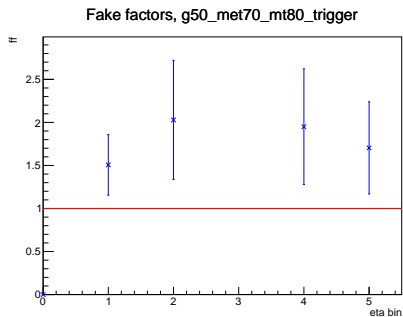
(a) Analysis Trigger,  $\eta$  inclusive region.



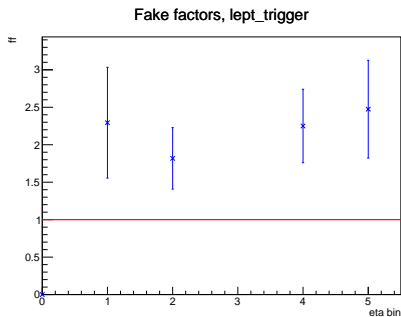
(b)  $p_T^{miss}$  Trigger,  $\eta$  inclusive region.

Figure 13: Fake factors stability check for  $p_T^{miss}$  Trigger in the  $\eta$  inclusive region (right) and Analysis Trigger in the  $\eta$  inclusive region (left) using nominal  $R, R_b$ .

# Fake factors in different configurations



(a) Analysis Trigger,  $p_T$  inclusive region.



(b) Leptonic Trigger,  $p_T$  inclusive region.

Figure 14: Fake factors stability check for Leptonic Trigger in the  $p_T$  inclusive region (right) and Analysis Trigger in the  $p_T$  inclusive region (left) using nominal  $R$ ,  $R_b$ .



# Uncertainties on fake factors

We can now calculate the uncertainties on these fake factors as:

$$\sigma_{ff} = \sqrt{\sigma_{R_b}^2 + \sigma_R^2}$$

where

$$\sigma_{R_b} = \left( \frac{ff(R_{nom}, R_{b,up}) - ff(R_{nom}, R_{b,down})}{2} \right)$$

$$\sigma_R = \left( \frac{ff(R_{up}, R_{nom}) - ff(R_{down}, R_{nom})}{2} \right)$$

# Fake factors in different configurations

$\eta$ bin	ff	$\sigma_{ff}$	%
1	1.51	0.35	23.4 %
2	2.03	0.69	34.1 %
4	1.95	0.67	34.4 %
5	1.70	0.54	31.4 %

$\eta$ bin	ff	$\sigma_{ff}$	%
1	2.29	1.48	64.4 %
2	1.82	0.82	45.1 %
4	2.25	0.98	43.6 %
5	2.47	1.30	52.7 %

Table 2: Fake factors in  $p_T$  incl region using Analysis(l) and Leptonic Trigger(r).

$p_T$ bin	ff	$\sigma_{ff}$	%
3	0.71	0.12	17.4 %
4	0.71	0.19	27.6 %
5	1.50	0.67	44.6 %
6	1.36	0.83	61.1 %
7	1.82	1.03	56.9 %
8	1.34	0.31	23.4 %

$p_T$ bin	ff	$\sigma_{ff}$	%
3	0.94	0.80	84.6 %
4	0.67	0.63	95.3 %
5	1.02	0.80	79.2 %
6	1.46	1.08	73.9 %
7	1.30	1.19	91.6 %
8	1.27	0.79	62.7 %

Table 3: Fake factors in the  $\eta$  incl region using Analysis(l) and  $p_T^{miss}$  Trigger(r).

# Applying fake factors to CR

We extrapolate the number of jets faking photons in SR  $N_{\text{ext}}^{\text{SR}}(ff)$  and compute its uncertainty from the fake factors uncertainty  $\sigma_N$

$$\sigma_N = \frac{N(ff + \sigma_{ff}) - N(ff - \sigma_{ff})}{2}$$

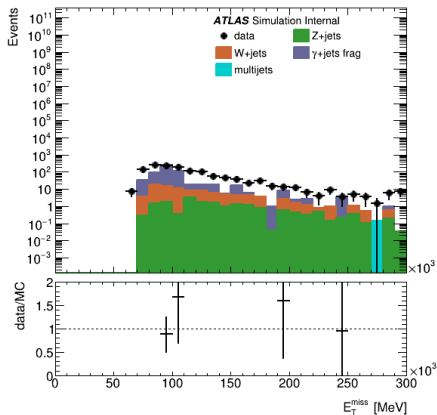
## Jet faking photons extrapolated with nominal/ varied fake factors

	$N_{\text{ext}}^{SR}(ff) \pm \sigma_N$	$N_{\text{ext}}^{SR}(ff - \sigma_{ff})$	$N_{\text{ext}}^{SR}(ff + \sigma_{ff})$
all	9.2e+06 $\pm$ 1.5e + 06	1.0745e+07	7.7145e+06
$n_e = 0$	9.2e+06 $\pm$ 1.5e + 06	1.0745e+07	7.7145e+06
$n_\mu = 0$	9.2e+06 $\pm$ 1.5e + 06	1.0745e+07	7.7145e+06
$p_T^{\text{miss}} > 100 \text{ GeV}$	1.4e+06 $\pm$ 2.2e + 05	1.6335e+06	1.1838e+06
$n_\gamma^{\text{nonisol}}$	2.7e+05 $\pm$ 4.2e + 04	3.1206e+05	2.2676e+05
$p_T^\gamma > 50 \text{ GeV}$	2.6e+05 $\pm$ 4.2e + 04	3.0792e+05	2.2377e+05
$n_{\text{jet}}$	2.1e+05 $\pm$ 3.3e + 04	2.4302e+05	1.7653e+05
$m_T > 80 \text{ GeV}$	2.1e+05 $\pm$ 3.3e + 04	2.4175e+05	1.756e+05
$\Delta\Phi(\vec{p}_T^{\text{miss}}, [\vec{p}_T^{\text{miss}}]_\gamma) \geq 1.25$	8146 $\pm$ 1291	9440	6857
$S_{p_T^{\text{miss}}} > 6$	1695 $\pm$ 267	1963	1429
$\Delta p_T^{\text{miss}} > -10 \text{ GeV}$	1311 $\pm$ 206	1518	1105
$ \eta_\gamma  \leq 1.75$	1022 $\pm$ 160	1183	863
$\Delta\Phi(\vec{p}_T^{\text{miss}}, [\vec{p}_T^{\text{miss}}]_j) \leq 0.75$	926 $\pm$ 145	1073	782
$\Delta\Phi(\vec{p}_T^{j1}, \vec{p}_T^{j2}) \leq 2.5$	775 $\pm$ 116	897	655

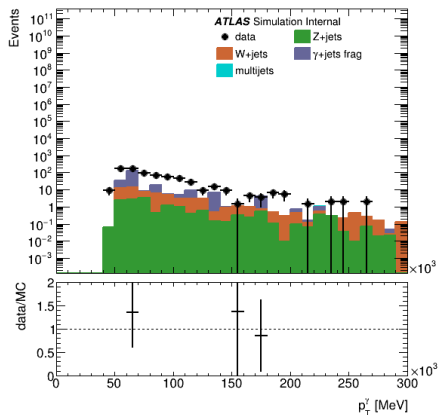
## Monte Carlo samples in Signal Region

	multijets	$\gamma$ +jets(f)	W+jets	Z+jets	sum
all	6.2709e+08	1.9313e+06	21836	13044	6.2906e+08
$n_{e\ell} = 0$	6.2709e+08	1.9313e+06	21836	13044	6.2906e+08
$n_{\mu} = 0$	6.2709e+08	1.9313e+06	21836	13044	6.2906e+08
$p_T^{miss} > 100$ GeV	3.691e+08	2.3888e+05	11572	8612	3.6936e+08
$n_{\gamma}^{isol} = 1$	5.3566e+07	1.1368e+05	1282	701.51	5.3682e+07
$p_T^{\gamma} > 50$ GeV	5.3566e+07	1.1248e+05	1276	699.92	5.3681e+07
$n_{jet} < 4$	3.8745e+05	82875	1112	630.72	4.7207e+05
$m_T > 80$ GeV	3.8736e+05	82259	1061	626.18	4.713e+05
$\Delta\Phi(\vec{p}_T^{miss}, [\vec{p}_T^{miss}]_{\gamma}) \geq 1.25$	2552.8	5230	133.89	43.494	7959.9
$S_{p_T^{miss}} > 6$	344.28	1001	104.31	36.287	1486.2
$\Delta p_T^{miss} > -10$ GeV	215.74	591.05	95.261	34.069	936.12
$ \eta_{\gamma}  < 1.75$	165.45	383.81	71.445	28.204	648.91
$\Delta\Phi(\vec{p}_T^{miss}, [\vec{p}_T^{miss}]_{jet}) \leq 0.75$	137.74	338.96	63.101	23.481	563.28
$\Delta\Phi(p_T^{j1}, p_T^{j2}) \leq 2.5$	0.6921	166.28	56.848	22.082	245.91

## Comparison between MC and data-driven estimation

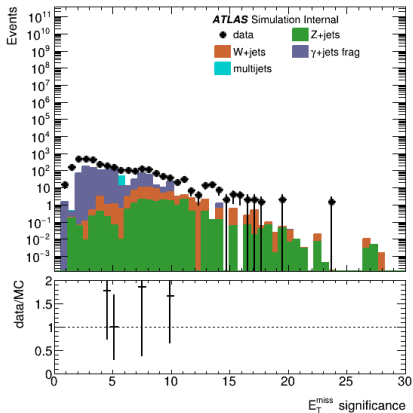


(a)  $p_T^{\text{miss}}$

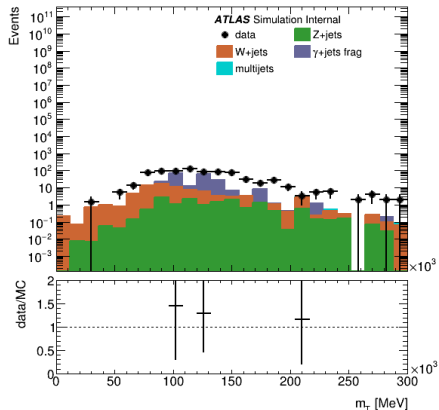


(b)  $p_T^\gamma$

## Comparison between MC and data-driven estimation

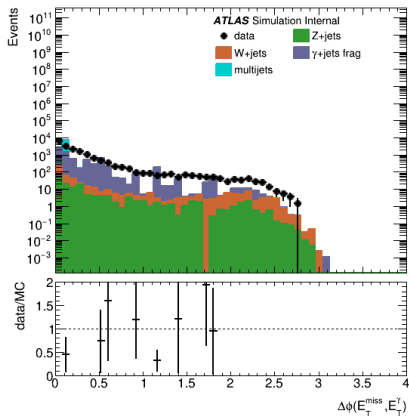


(a)  $S_{\rho_T^{\text{miss}}}$

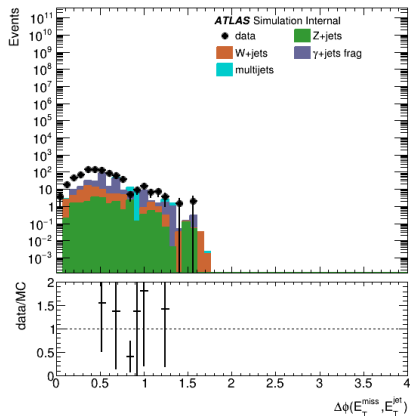


(b)  $m_T$

## Comparison between MC and data-driven estimation



(a)  $\Delta\Phi(\vec{p}_T^{\text{miss}}, [\vec{p}_T^{\text{miss}}]_\gamma)$



(b)  $\Delta\Phi(\vec{p}_T^{\text{miss}}, [\vec{p}_T^{\text{miss}}]_{\text{jet}})$

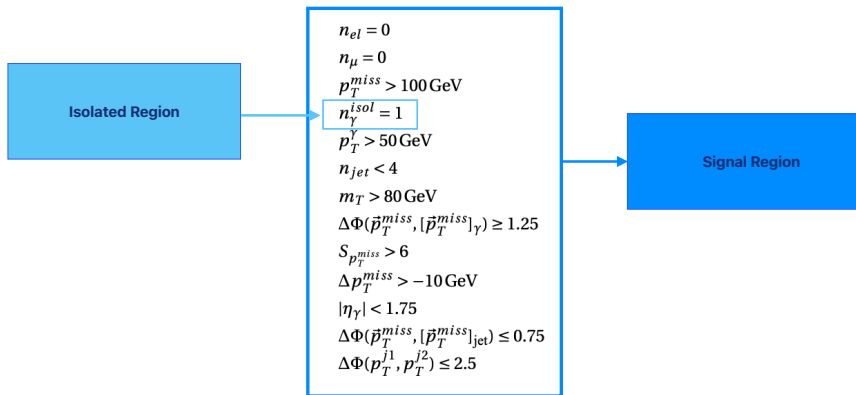


# Non-Isolated Control Region

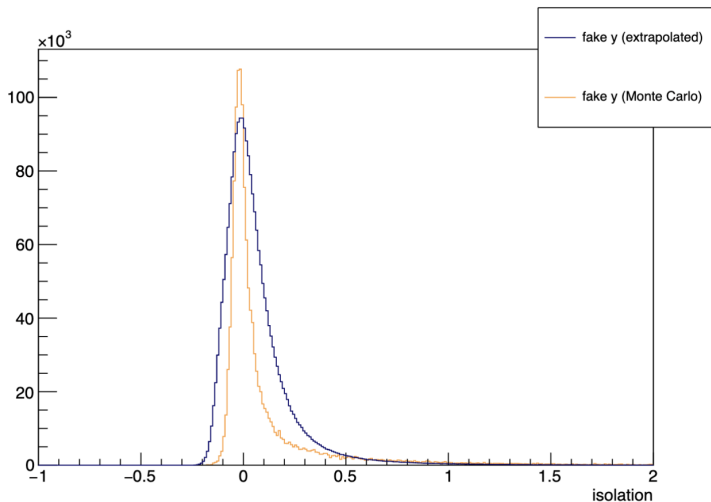
Non-Isolated Region

$$\begin{aligned}n_{el} &= 0 \\n_{\mu} &= 0 \\p_T^{miss} &> 100 \text{ GeV} \\n_{\gamma}^{nonisol} &= 1 \\p_T^{\gamma} &> 50 \text{ GeV} \\n_{jet} &< 4 \\m_T &> 80 \text{ GeV} \\ \Delta\Phi(\vec{p}_T^{miss}, [\vec{p}_T^{miss}]_{\gamma}) &\geq 1.25 \\S_{p_T^{miss}} &> 6 \\ \Delta p_T^{miss} &> -10 \text{ GeV} \\|\eta_{\gamma}| &\leq 1.75 \\ \Delta\Phi(\vec{p}_T^{miss}, [\vec{p}_T^{miss}]_{jet}) &\leq 0.75 \\ \Delta\Phi(p_T^{j1}, p_T^{j2}) &\leq 2.5\end{aligned}$$

Control Region









# Comparison between extrapolated isolation distribution and MC isolation distribution



$W\gamma$

## Signal Region

- $N_\gamma^{isol} = 1$ ;
- $N_e = 0$ ;
- $N_\mu = 0$ ; 
- $|\vec{p}_T^{miss}| > 100$  GeV; 
- $|\vec{p}_T^\gamma| > 50$  GeV;
- $N_{jets} \leq 3$ ;
- $m_T > 80$  GeV; 
- $\Delta\Phi(\vec{p}_T^{miss}, [\vec{p}_T^{miss}]_\gamma) \geq 1.25$ ; 
- $S_{p_T^{miss}} > 6$ ;
- $\Delta|\vec{p}_T^{miss}| > -10$  GeV; 
- $\Delta\Phi(\vec{p}_T^{miss}, [\vec{p}_T^{miss}]_{jet}) \leq 0.75$ ; 
- $\Delta\Phi(\vec{p}_T^{j1}, \vec{p}_T^{j2}) \leq 2.5$ .

## $1\mu$ Control Region

- $N_\gamma^{isol} = 1$ ;
- $N_e = 0$ ;
- $N_\mu = 1$ ;
- $|[\vec{p}_T^{miss}]_{no\ \mu}| > 100$  GeV;
- $|\vec{p}_T^\gamma| > 50$  GeV;
- $N_{jets} \leq 3$ ;
- $[m_T]_{no\ \mu} > 80$  GeV;
- $\Delta\Phi([\vec{p}_T^{miss}]_{no\ \mu}, [\vec{p}_T^{miss}]_\gamma) \geq 1.25$ ;
- $S_{[p_T^{miss}]_{no\ \mu}} > 6$ ;
- $\Delta|[\vec{p}_T^{miss}]_{no\ \mu}| > -10$  GeV;
- $\Delta\Phi([\vec{p}_T^{miss}]_{no\ \mu}, [\vec{p}_T^{miss}]_{jet}) \leq 0.75$ ;
- $\Delta\Phi(\vec{p}_T^{j1}, \vec{p}_T^{j2}) \leq 2.5$ .

We would like to calculate a K-factor as ratio between the number of events in data and Monte Carlo in the Control Region, to **correct** the imperfections of the simulations.

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The K-factor will be later applied to Monte-Carlo in the Signal Region to get the **final estimation** of the background yield in Signal Region.

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The K-factor will be later applied to Monte-Carlo in the Signal Region to get the **final estimation** of the background yield in Signal Region.

We calculated K-factors in 3 **different ways**:

- for all the Monte Carlo samples (approach 1);
- for  $W\gamma + W\text{jets}$  sample (approach 2);
- for  $W\gamma$  sample only, using a data-driven estimation of  $W\text{jets}$  events (approach 3).  
→ Fake factors computed as described in the last meeting (link<sup>1</sup>) were used!

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<sup>1</sup><https://indico.cern.ch/event/1420907/contributions/5975099/attachments/2864837/5013945/JetsFakingPhotons.pdf>

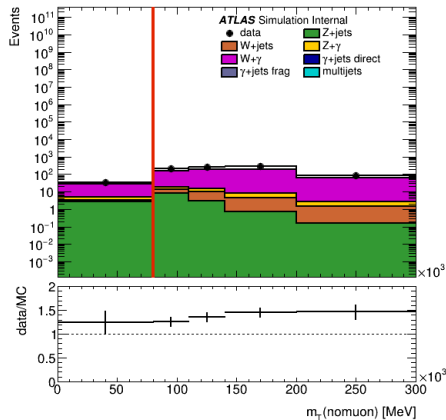


K-factors are calculated in bins of transverse mass  $m_T$ , a good candidate to be a discriminating variable.

- bin 1:  $[m_T]_{no\mu} \in [0, 80]\text{GeV} \rightarrow$  cut on  $m_T$  defining the CR;
- bin 2:  $[m_T]_{no\mu} \in [80, 110]\text{GeV}$ ;
- bin 3:  $[m_T]_{no\mu} \in [110, 140]\text{GeV} \rightarrow$  centred on  $m_H$ , expected for signal;
- bin 4:  $[m_T]_{no\mu} \in [140, 200]\text{GeV}$ ;
- bin 5:  $[m_T]_{no\mu} > 200 \text{ GeV}$ .

# Approach 1: K-factors as data/MC

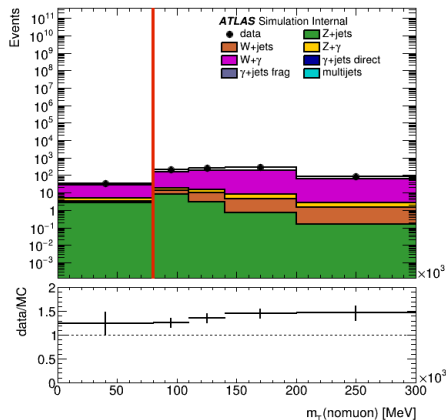
$$K(m_T) = \frac{N_{data}^{1\mu CR}(m_T)}{N_{MC}^{1\mu CR}(m_T)}$$



# Approach 1: K-factors as data/MC

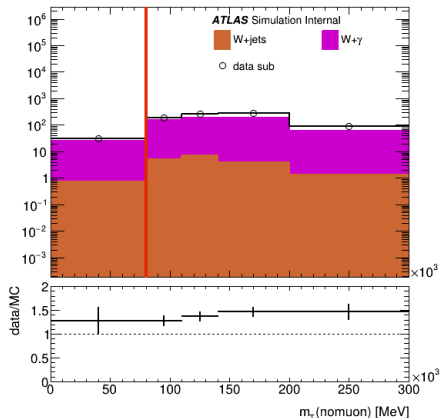
$$K(m_T) = \frac{N_{data}^{1\mu CR}(m_T)}{N_{MC}^{1\mu CR}(m_T)}$$

$m_T$ (GeV)	K	$\sigma_K^{stat}$
0-80	1.240	0.244
80-110	1.259	0.100
110-140	1.361	0.097
140-200	1.459	0.096
>200	1.465	0.156



# Approach 2: K-factors for $W\gamma+W$ jets

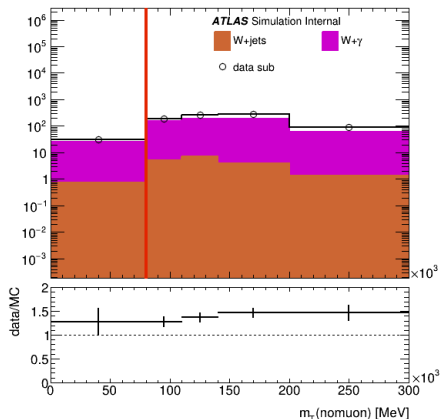
$$K(m_T) = \frac{[N_{data}^{1\mu CR} - N_{Zjets}^{1\mu CR} - N_{Z\gamma}^{1\mu CR} - N_{\gamma jets}^{1\mu CR} - N_{multijets}^{1\mu CR}](m_T)}{[N_{W\gamma}^{1\mu CR} + N_{Wjets}^{1\mu CR}](m_T)}$$



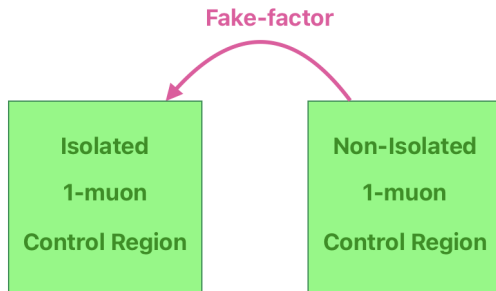
# Approach 2: K-factors for $W\gamma+W\text{jets}$

$$K(m_T) = \frac{[N_{data}^{1\mu CR} - N_{Z\text{jets}}^{1\mu CR} - N_{Z\gamma}^{1\mu CR} - N_{\gamma\text{jets}}^{1\mu CR} - N_{\text{multijets}}^{1\mu CR}](m_T)}{[N_{W\gamma}^{1\mu CR} + N_{W\text{jets}}^{1\mu CR}](m_T)}$$

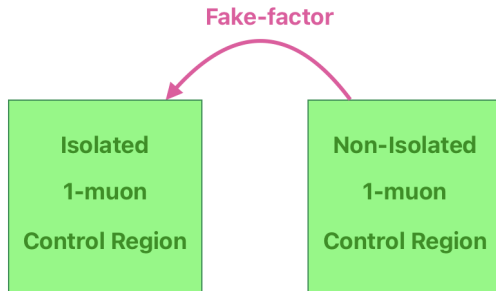
$m_T$ (GeV)	K	$\sigma_K^{stat}$
0-80	1.279	0.287
80-110	1.282	0.109
110-140	1.376	0.101
140-200	1.473	0.099
>200	1.476	0.160



# Jets faking photons in the $1 \mu$ Control Region



# Jets faking photons in the $1 \mu$ Control Region



$$N_{dd}^{1\mu CR} = 276 \pm 84$$

# Applying K-factors to MC in Signal Region

cut	$N_{W\gamma}^{SR}(K)$	$N_{W\gamma}^{SR}(K + \sigma_K)$	$N_{W\gamma}^{SR}(K - \sigma_K)$
all	28543	22835	34250
$n_e = 0$	28543	22835	34250
$n_\mu = 0$	28543	22835	34250
$n_\tau = 0$	23343	18681	28003
$p_T^{miss} > 100 \text{ GeV}$	10143	8062.1	12223
$n_\gamma^{isol}$	7408.6	5882.7	8934.3
$p_T^\gamma > 50 \text{ GeV}$	7387.8	5866.1	8909.5
$n_{jet}$	6332.2	5029.4	7635.1
$m_T > 80 \text{ GeV}$	6254.6	4972.9	7536.4
$\Delta\Phi(\vec{p}_T^{miss}, [\vec{p}_T^{miss}]_\gamma) \geq 1.25$	822.81	655.64	989.99
$S_{p_T^{miss}} > 6$	711.51	566.92	856.09
$\Delta p_T^{miss} > -10 \text{ GeV}$	650.97	518.67	783.28
$ \eta_\gamma  \leq 1.75$	490.74	391.01	590.47
$\Delta\Phi(\vec{p}_T^{miss}, [\vec{p}_T^{miss}]_j) \leq 0.75$	425.12	338.69	511.55
$\Delta\Phi(\vec{p}_T^{j1}, \vec{p}_T^{j2}) \leq 2.5$	369.38	294.26	444.5